

Coherent and incoherent band dispersions in V 3d band of SrVO<sub>3</sub> thin filmsMasaru TAKIZAWA\*<sup>1</sup>, Makoto MINOHARA<sup>2</sup>, Hiroshi KUMIGASHIRA<sup>3,4</sup>, Akira CHIKAMATSU<sup>3</sup>, Masaharu OSHIMA<sup>2,3,4</sup>, Teppei YOSHIDA<sup>1</sup>, and Atsushi FUJIMORI<sup>1</sup><sup>1</sup>Department of Physics, University of Tokyo, Bunkyo-ku, Tokyo 113-0033, Japan<sup>2</sup>Graduate School of Arts and Sciences, University of Tokyo, Meguro-ku, Tokyo 153-8902, Japan<sup>3</sup>Department of Applied Chemistry, University of Tokyo, Tokyo 113-8656, Japan<sup>4</sup>Research for Evolutional Science and Technology of Japan Science and Technology Corporation (JSTCREST), Chiyoda-ku, Tokyo 102-0075, Japan**Introduction**

Metal-insulator transition has been extensively studied because of its fundamental importance as well as its close relationship to interesting phenomena such as high-temperature superconductivity in cuprates and colossal magnetoresistance in manganites [1]. Ca<sub>1-x</sub>Sr<sub>x</sub>VO<sub>3</sub> (CSVO) is a typical bandwidth control system but remains metallic in the entire  $x$  range. With Ca doping, ultra-violet photoemission spectra of CSVO showed spectral weight transfer from the coherent part to the incoherent part [2], while using high photon energies there were no spectral weight transfer [3]. Therefore, it is now well known that the surface electronic states are different from the bulk ones. Many studies were devoted to investigate the real “bulk” electronic states [4-6], but the problem remains highly controversial and further studies are strongly required. In the present work, we have fabricated a SrVO<sub>3</sub> (SVO) thin film and studied its electronic structure in detail by angle-resolved photoemission spectroscopy (ARPES).

**Experiment**

A SVO thin film was fabricated in a laser MBE chamber connected to a synchrotron radiation ARPES system at BL-28B of Photon Factory [7]. The films were deposited on Nb-doped TiO<sub>2</sub>-terminated SrTiO<sub>3</sub> (001) substrates [8] at 900 °C at an ultra high vacuum of  $\sim 10^{-9}$  Torr. The fabricated SVO thin film was transferred into the photoemission chamber under an ultrahigh vacuum of  $10^{-10}$  Torr. The ARPES spectra were taken at 20 K with the total energy resolution of 30 meV.

**Results and Discussion**

The  $E - k_x$  space intensity plot near  $E_F$  along the  $\Gamma - X$  direction in Fig. 1 shows the V 3d bands. The peak positions determined from both energy distribution curves (EDCs) and momentum distribution curves (MDCs) are also shown. The V 3d<sub>xy</sub> and 3d<sub>xz</sub> bands cross the Fermi level between the  $\Gamma$  and X points. For the dispersion of the coherent part, one can see clear mass renormalization compared with the LDA calculation [9]. From the experimental  $(-0.44 \pm 0.02$  eV) and calculated  $(-0.95$  eV) occupied bandwidths, the global mass renormalization factor is estimated to be  $\sim 2$ . That is, if the LDA band dispersions are reduced by a factor of 0.5, the experimental band dispersions are well reproduced as shown in Fig. 1. This indicates that the self-energy is nearly independent of momentum and of the  $d_{xy}$ ,  $d_{yz}$  or  $d_{xz}$

bands of the degenerate  $t_{2g}$  band. The kink in the band dispersion is weak and broad, if exists, but the curvature changes its sign around  $\sim -0.2$  eV as predicted by a recent DMFT calculation [10]. As for the incoherent part located around  $-1.5$  eV, one can see a weak but finite ( $\sim 0.1$  eV) dispersion. The intensity of the incoherent part is momentum dependent and becomes strong within the Fermi surface.

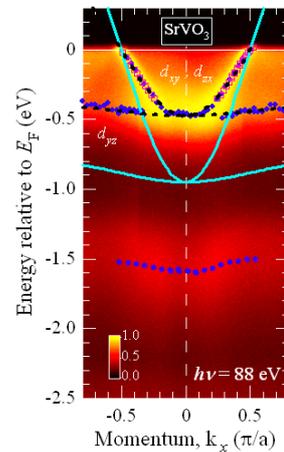


Figure 1: Energy- and momentum-dependent spectral weight near the Fermi level for SrVO<sub>3</sub>. Peak positions of the EDCs and MDCs are shown by filled circles and open squares, respectively. The V 3d bands from the LDA calculation [9] is also shown by solid curves. Broken curves are LDA bands renormalized by a factor of 2.

**References**

- [1] M. Imada, A. Fujimori, and Y. Tokura, Rev. Mod. Phys. **70** 1039 (1998), and references therein.
- [2] I. H. Inoue *et al.*, Phys. Rev. Lett. **74**, 2539 (1995).
- [3] A. Sekiyama *et al.*, Phys. Rev. Lett. **93**, 156402 (2004).
- [4] R. Eguchi *et al.*, Phys. Rev. Lett. **96**, 076402 (2006).
- [5] K. Maiti *et al.*, Phys. Rev. B **73**, 052508 (2006).
- [6] T. Yoshida *et al.*, Phys. Rev. Lett. **95**, 146404 (2006).
- [7] K. Horiba *et al.*, Rev. Sci. Instr. **74**, 3406 (2003).
- [8] M. Kawasaki *et al.*, Science **266**, 1540 (1994).
- [9] E. Pavarini *et al.*, New J. Phys. **7**, 188 (2005).
- [10] I. A. Nekrasov *et al.*, Phys. Rev. B **73**, 155112 (2006).

\* takizawa@wyvern.phys.s.u-tokyo.ac.jp